Coefficient of Thermal expansion

Since we have just discussed heat capacity at constant volume, it is worth discussing the closely related coefficient of thermal expansion. When objects are heated, they expand. The fractional change in the volume of a material expands as its temperature increases is called the

$$\alpha_{V} = \frac{1}{V} \frac{\partial V}{\partial T} = \frac{\partial \ln(V)}{\partial T}$$

Note that the expansion in any single dimension can be found from

$$\alpha_{V} = \frac{1}{V} \frac{\partial V}{\partial T} = \frac{1}{L^{3}} \frac{\partial L^{3}}{\partial T} = \frac{3L^{2}}{L^{3}} \frac{\partial L}{\partial T} = \frac{3}{L} \frac{\partial L}{\partial T}$$

So if we define a linear coefficient of thermal expansion, α_L as

$$\alpha_L = \frac{1}{L} \frac{\partial L}{\partial T}$$

then $\alpha_V = 3\alpha_L$.

Coefficient of thermal expansion written in terms of density

$$\alpha_{V} = \frac{1}{V} \frac{\partial(V)}{\partial T} = \frac{1}{\frac{1}{\rho}} \frac{\partial\left(\frac{1}{\rho}\right)}{\partial T} = \rho \frac{\partial(\rho^{-1})}{\partial T} = -\frac{\rho}{\rho^{2}} \frac{\partial(\rho)}{\partial T} = -\frac{1}{\rho} \frac{\partial(\rho)}{\partial T} = -\frac{\partial\ln(\rho)}{\partial T}$$

Coefficient of thermal expansion of a gas

For an ideal gas we have $P = n R^* T = \rho T R^*/m$. So $\rho = m P/(T R^*)$. If we hold pressure constant, then the volume coefficient of thermal expansion of an ideal gas is

$$\alpha_{V} = -\frac{1}{\rho} \frac{\partial(\rho)}{\partial T} = -\frac{R * T}{mP} \frac{mP}{R *} \frac{\partial(T^{-1})}{\partial T} = +\frac{T}{T^{2}} \frac{\partial(T)}{\partial T} = \frac{1}{T}$$

Examples of the volume coefficient of thermal expansion

| In units of ppm/°C | |
|--------------------|------|
| Water | 207 |
| Mercury | 182 |
| Ethanol | 750 |
| Glass | 25.5 |

Mercury and alcohol thermometers

Mercury and alcohol thermometers are based on thermal expansion of these liquids. The idea is to use a large reservoir of the liquid and force the thermal expansion into a small volume with a small cross section such that its height change is maximum.

The initial volume is V. So the increase (or decrease) in volume is $\alpha_v V \Delta T$. The thermometer design is such that this volume increase moves into a chamber with a small crossectional area, A, such that the height of the liquid in the chamber, ΔL , is amplified.

$$A \Delta L = \alpha_v V \Delta T$$
$$\Delta L = \frac{\alpha_v V \Delta T}{A}$$

So to make ΔL as large as possible, we want to use a liquid with as large thermal expansion coefficient as possible as well as making the *V*/*A* ratio as large as possible meaning as large a reservoir as possible and as small a crosssection as possible.

The advantage of ethanol is it has a very large thermal expansion coefficient of 750 ppm/K. Mercury has a different advantage. In the volume above the liquid, there is in theory a vacuum. In reality, a small number of gas molecules exists in the volume. This causes a pressure pushing down on the liquid surface which suppresses a bit of the thermal expansion. The advantage of mercury is its very low saturation vapor pressure which minimizes the pressure pushing down on the mercury liquid surface.

There is a combination that is particularly effective which is to put alcohol in the volume and put a small amount of mercury above the alcohol in the narrow tube. This takes advantage of the large thermal expansion coefficient of alcohol *AND* the small vapor pressure of mercury.

Ocean Warming and Sea Level Rise



As the ocean warms, the ocean expands. The amount of sea level rise depends on the temperature increase and the amount or percentage of ocean water that is warming. The ocean is heated from above and is therefore stable. Therefore it will take a long time for the entire depth of the ocean to warm. The portion of the ocean that is warmed from above is called the thermocline. The direct solar heating is much shallower than 500 m but currents and overturning

mix the warming through the depth of the thermocline. We assume the depth that experienced the enhanced warming is the depth of the thermocline.

The increase in greenhouse gas concentrations has caused reduced the IR emission to space. It has also increased the IR emission from the atmosphere into the surface. This will cause the surface to warm. Some of this absorbed energy will be stored in the surface and some will be released from the surface into the atmosphere and space by increased evaporation, increased conductive transfer to the atmospheric molecules in contact with the ocean surface and increased IR radiation emitted by the surface.



The net amount of heating that remains in the surface is difficult to estimate. We can simply use measurements of the rise in ocean temperature to determine the net heating of the ocean. We can also work backwards to estimate the net heat flowing into the oceans.



The figure above indicates SST has increased by ~0.5 K over the past 40 years which is a rate of ~0.0125 K/year.

The NET energy flux into the ocean

We know the measured change in the ocean surface temperature. To calculate the corresponding net (In minus Out) energy flux into the ocean we need to know how much energy storage capacity is associated with this rise in temperature.

The heat capacity of liquid water is 4,181 J/kg/K. We assume the depth of the heating is about 500 m to correspond with the thermocline. So a column of heated liquid water with cross sectional area of 1 m^2 and a depth of 500 m has a volume of 500 m³. The mass of that water column with a density of 1000 kg/m³ is $5x10^5$ kg. The rate of increase in the internal energy of the column is related to the rate of change of the temperature of the column as

$$\frac{dU}{dt} = mC\frac{dT}{dt}$$

where m is the mass of the column and C is the heat capacity or specific heat of liquid water. The answer is

$$5x10^{5}$$
 kg $4.2x10^{3}$ J/kg/K $1.25x10^{-2}$ K/year = 0.8 W

The 4 W/m^2 we used in the homework was the *decrease* in IR to space resulting from an instantaneous doubling of CO₂. This 0.8 W/m^2 flux is the downward *increase* in IR minus the increase in the fluxes cooling the ocean due to dry convection, evaporation, and radiative fluxes at the top of the ocean.

Thermal expansion of the ocean

Assuming the top 500 m of the ocean surface is warming at a rate of 0.0125 K/year suggests that that layer should be expanding and the sea level rising. Assume that the layer of ocean that is warming has an area, A, and a depth, L, and a volume,

$$V = AL$$

Therefore the change in volume due to thermal expansion is

$$\Delta V = A \Delta L$$

Therefore the fractional change in volume is

$$\frac{\Delta V}{V} = \frac{A\Delta L}{AL} = \frac{\Delta L}{L}$$

The resulting change in sea level is

$$\Delta L = L \, \Delta V / V = L \, \alpha_v \, \Delta T$$

and the rate of change of sea level is

$$\frac{\Delta L}{\Delta t} = L\alpha_v \frac{\Delta T}{\Delta t}$$

Plugging in values, we get that the change in the height of the ocean due to thermal expansion of the top 500 m of the ocean is approximately





The measured rise in sea level is about 20 cm = 200 mm in 100 years or about 2 mm/year. If the depth to which the warming extends were greater than 500 m, then our calculated rate of sea level rise would be higher and closer to the observed rate. Alternatively, a factor contributing to the discrepancy is the fact that the observed rate also includes the contribution of land ice melt. Direct estimates of land ice melt are VERY uncertain.



The figure above shows the sea level rise measured by the TOPEX and JASON satellites. The average rate is about 37 mm in 12.5 years or about 3 mm/year. This is about 50% faster than the non-satellite estimate over the past 100 years. Note that in the previous figure, the slope of the satellite data is indeed larger than the tide gauge-based slope.



Trend of Sea Level Change (1993-2008)

The figure above shows the regional change in sea level observed by altimetry satellites. Note how inhomogeneous this is. This is in large part due to ocean current. Warm water is piling up in the tropical western Pacific because the easterly low latitude winds push the surface water there. Water in the tropical eastern Pacific remains cooler because of upwelling water from depth there. The water at depth is cooler because it is below the thermocline and has not been affected by solar and IR radiation for a long time.